Methods for Determining Stellar Properties

Distance

- **❖** Parallax
 - \triangleright Distances less than ~750 light years
- ❖ Standard-Candle Method
 - > Uses apparent brightness and known luminosity

Temperature

- ❖ Wein's Law
 - > Color-temperature relation
- Spectral Class
 - ➤ O hot, M cool

Luminosity

- Inverse-Square Law
 - > Uses apparent brightness and known distance
- Luminosity Class
 - > I, II, III, IV, V

Composition

- Spectra
 - ➤ Observe known lines

Radius

- ❖ Stefan-Boltzmann Law
 - ➤ Measure luminosity & temperature
 - > Solve for radius
- **❖** Interferometer
 - > Triangulation
 - Uses angular size & distance
- Eclipsing Binary
 - ➤ Light curve

Mass

- **❖** Binary Stars
 - ➤ Modified form of Kepler's third law
- Mass-Luminosity Relation
 - > Brighter main sequence stars are more massive

Radial Velocity

- Doppler Shift
 - > Shifted spectral lines